

## **Aura Validation : H<sub>2</sub>O and N<sub>2</sub>O Subgroup Session**

### **TES**

Bob Herman

### **MLS**

Alyn Lambert (N<sub>2</sub>O and stratospheric/mesospheric H<sub>2</sub>O)

Bill Read (tropospheric H<sub>2</sub>O and RH)

### **HIRDLS**

John Gille

### ***In situ/In situ and In situ/MLS comparisons***

Elliot Weinstock

Holger Vömel

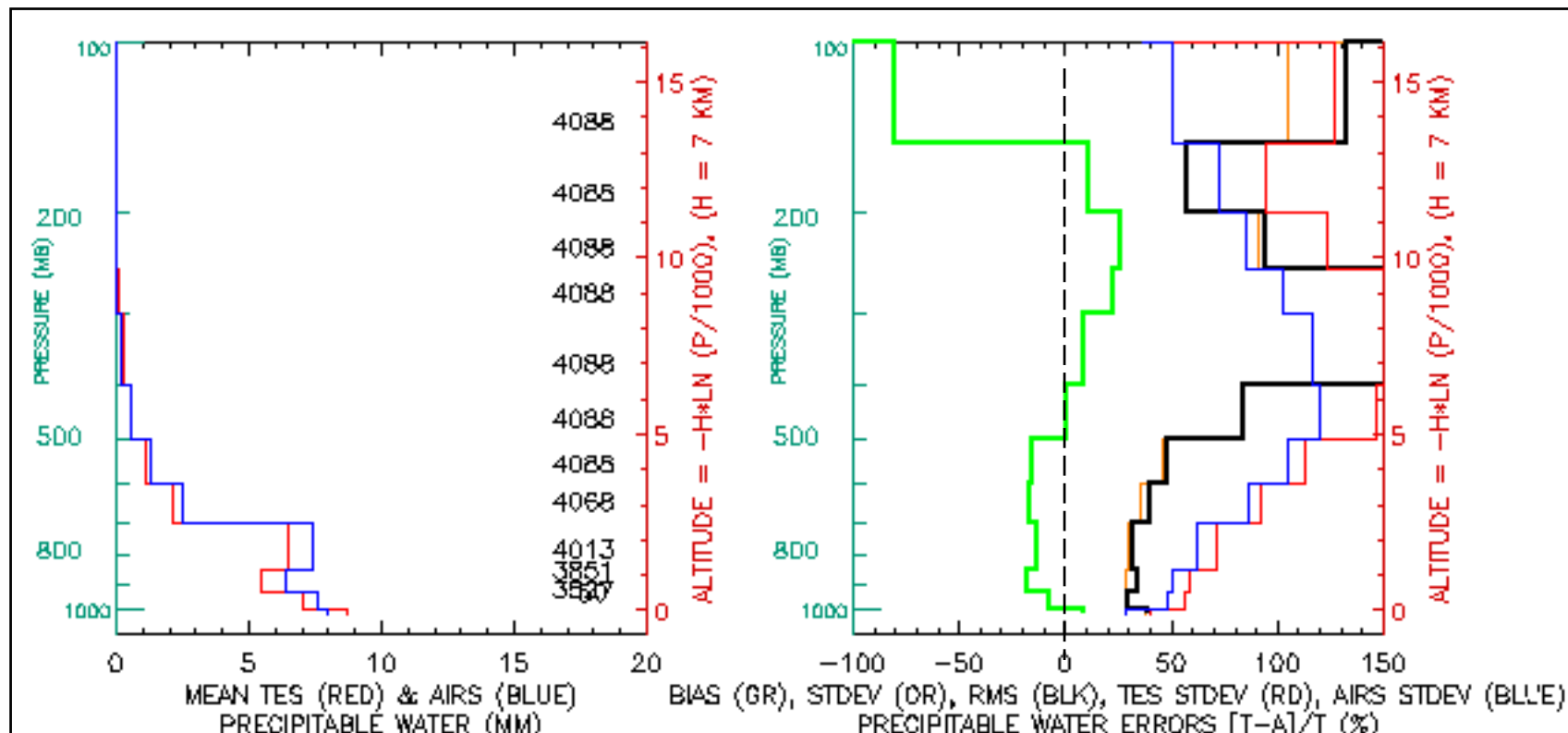
### **General Discussion**

TES v002 water comparisons (this is the version currently on the DAAC)

TES 15-20% drier than AIRS at 500-1000 hPa.

TES 10-25% wetter than AIRS at 150-500 hPa.

*Bob Herman*



Mean profiles

TES - AIRS

% Bias in green ( $[TES-AIRS]/TES$ ),  
rms differences in black

Note: Little latitudinal dependence on TES/AIRS differences.

# TES compared with sondes, averaging kernel applied to sonde data prior to comparisons.

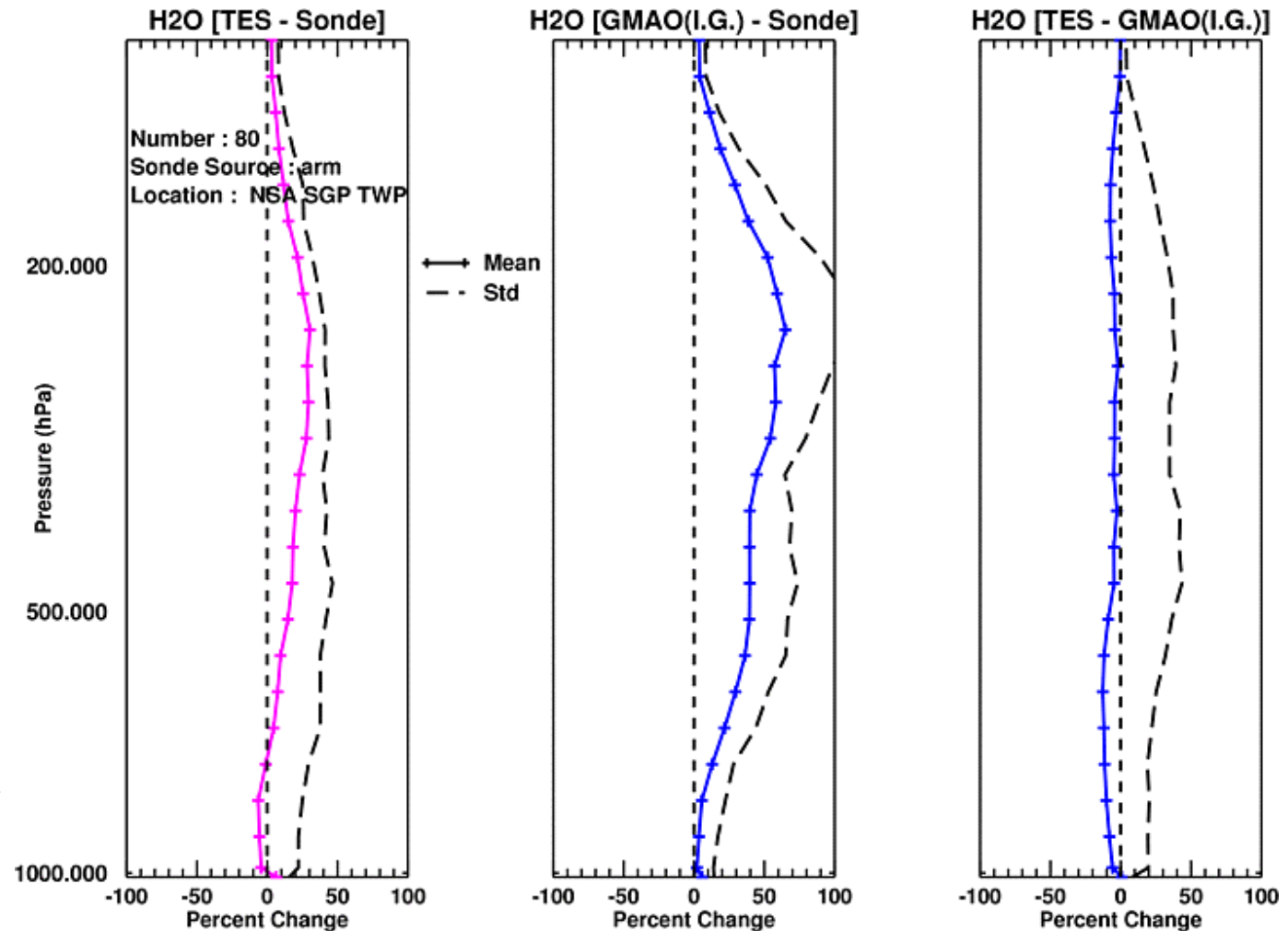
*Bob Herman*

80 radiosondes (RS90 and RS92) compared with TES special obs. at DOE ARM sites.

Coincidence criteria:

within 2 hours and 250 km of the sonde launch.

TES 0-30% wetter than sondes at 100-700 hPa.



Sonde comparisons demonstrate that TES improves on GMAO H<sub>2</sub>O

## TES SUMMARY

- TES v002 is 10-25% wetter than AIRS at 150-500 hPa.
- TES v002 is 15-20% drier than AIRS at 500-1000 hPa.
- TES v002 is 0-30% wetter than ARM site radiosondes at 100-700 hPa.
- The next release of TES data (v003) is coming.
- Next step: a more thorough analysis of CFH, NCEP sondes and aircraft data (including INTEx).
- Future validation needs: TES limb water vapor and high-latitude measurements poleward of 50 degrees.

*Bob Herman*

# Comparison of MLS v2.10 N<sub>2</sub>O with v1.51

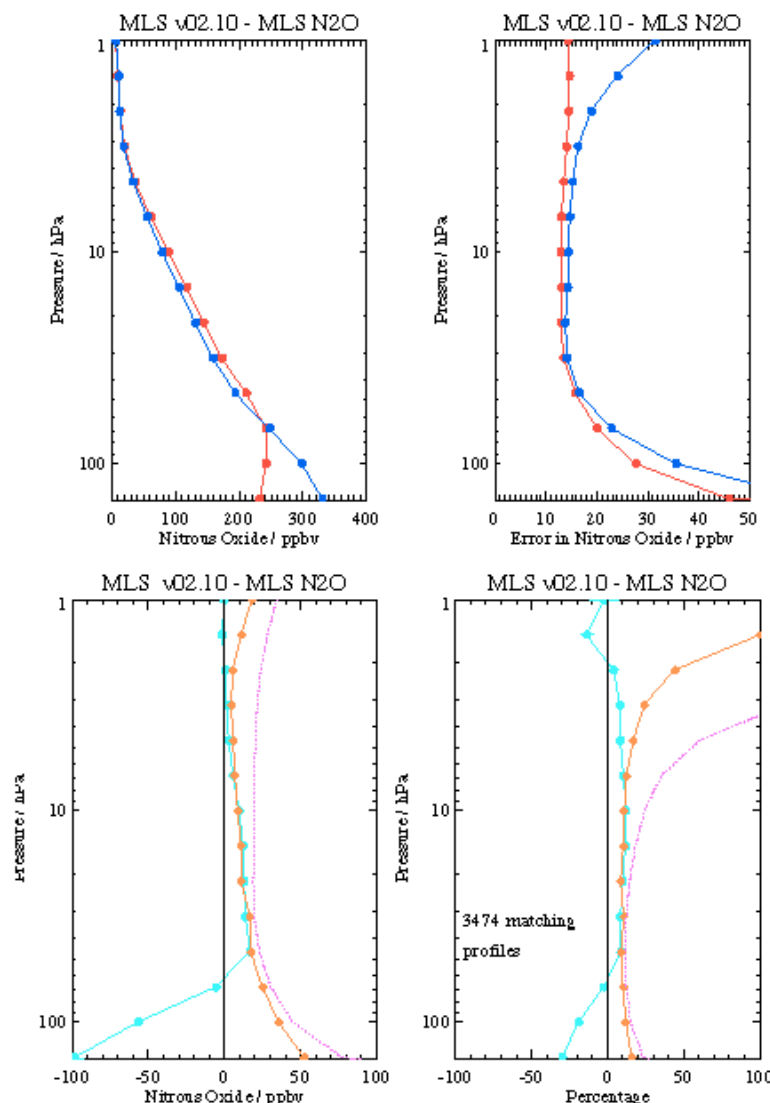
28 January 2005

- ▶ Upper panel: Global mean vertical profiles
  - ▶ MLS versions **v2.10** and **v1.51**
  - ▶ Left: Mean vmr profile
  - ▶ Right: RMS precision profile

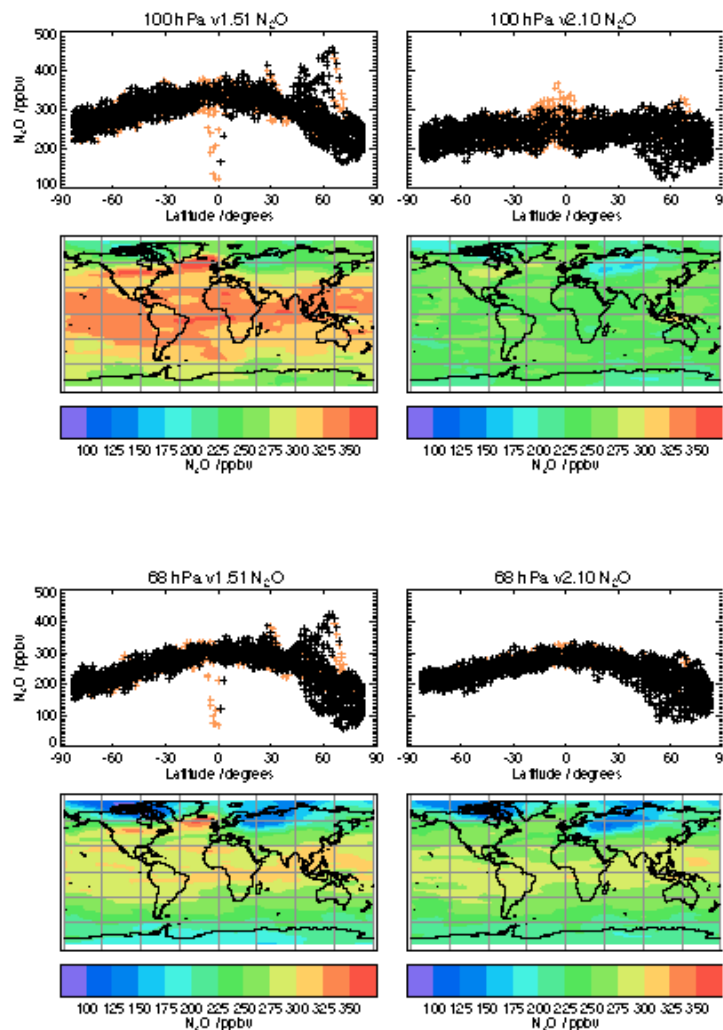
- ▶ Lower panel: Global mean difference vertical profiles

- ▶ Left: Statistics of profile differences in vmr units
  - ▶ Mean difference (**v2.10** – **v1.51**)
  - ▶ RMS difference
  - ▶ Expected RMS difference
- ▶ Right: as Left except shown as a percentage

*Alyn Lambert, N<sub>2</sub>O*



# Anomalies in the N<sub>2</sub>O Data Product



- ▶ Upper panel: 100 hPa N<sub>2</sub>O
  - ▶ Left: v1.51
    - ▶ Scatter plot: **orange points** indicate data marked by status/quality flags
    - ▶ Map: latitude – longitude distribution
  - ▶ Right: as Left except for v2.10
- ▶ Lower panel: as Upper panel except for 68 hPa N<sub>2</sub>O

Not all anomalous data points were detected by the status/quality flags in v1.51 (an off-line data mask is available)

v2.10 N<sub>2</sub>O is clearly too low at 100 hPa, but at 68 hPa the artifacts due to poor convergence in v1.51 have been improved

*Alyn Lambert, N<sub>2</sub>O*

# Example: ACE/MLS comparison

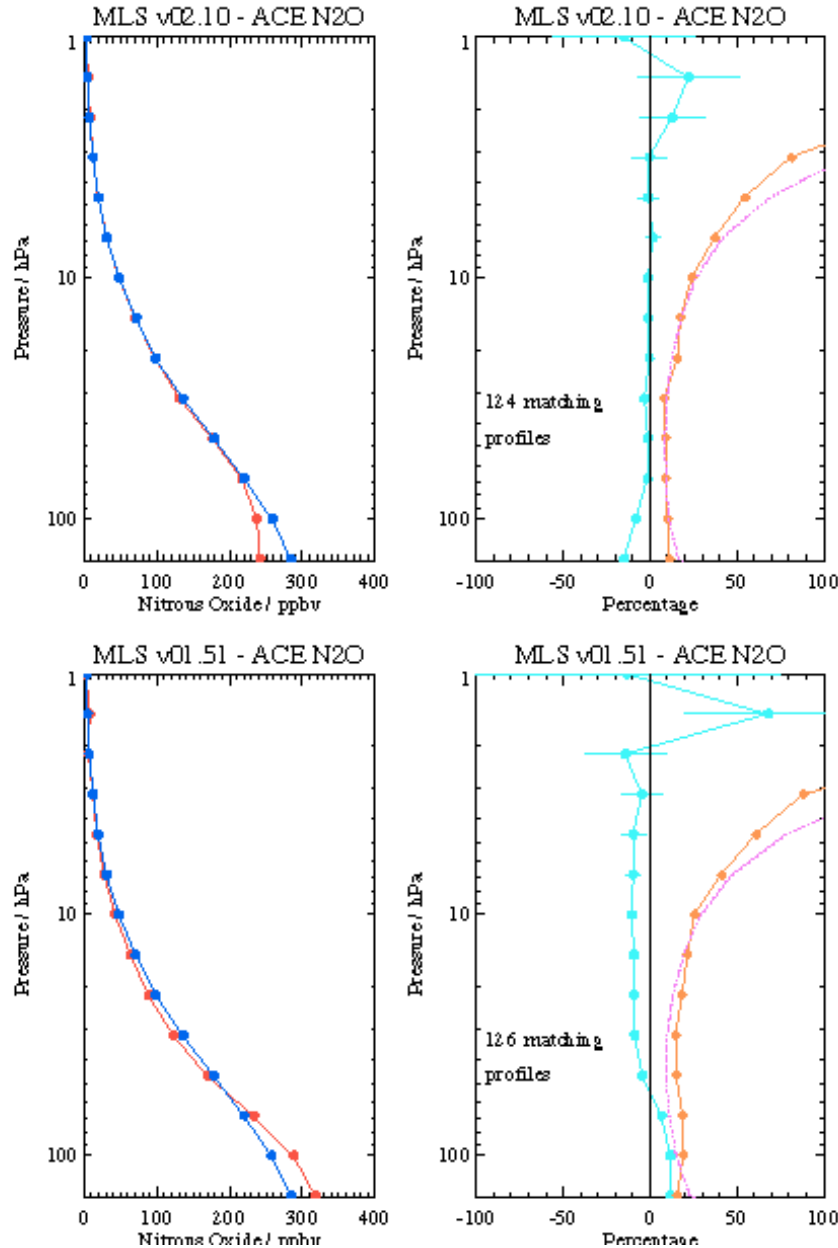
## N<sub>2</sub>O Summary

MLS v2.10 N<sub>2</sub>O is ~10% larger than v1.51 in the mid-stratosphere

In general, the biases and rms scatter of MLS v2.10 N<sub>2</sub>O against ACE, MIPAS and SMR are very good and show significant improvements over the MLS v1.51 data

Problems with poor convergence have been reduced in the MLS v2.10 retrievals

Further refinements will address the problem with low values of N<sub>2</sub>O at 100 hPa and greater pressures



*Alyn Lambert, N<sub>2</sub>O*

# Improvements to the retrieval scheme for v2.10

- ▶ Updated spectroscopy for H<sub>2</sub>O and contaminating species (O<sub>3</sub>, HNO<sub>3</sub>)
- ▶ Increased vertical resolution in the UTLS (12 levels per decade for pressures  $\geq 22$  hPa)
- ▶ Reduced error inflation (artificial increase of radiance noise to allow for uncertainties in the forward model)
- ▶ Included fine spectral resolution DACS radiances to improve mesospheric H<sub>2</sub>O retrievals

*Stratospheric water, MLS*

*Alyn Lambert, H<sub>2</sub>O*



# Comparison of MLS v2.10 H<sub>2</sub>O with v1.51

28 January 2005

v2.10 H<sub>2</sub>O is ~0.5 ppmv larger than v1.51

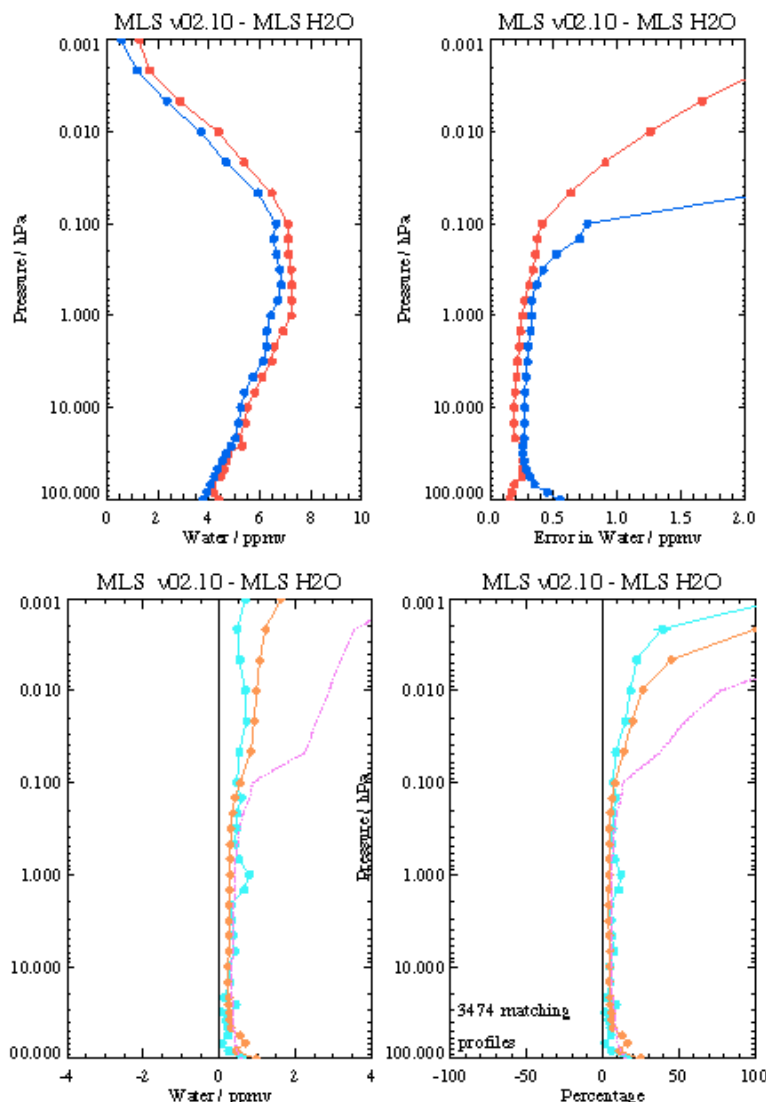
► Upper panel: Global mean vertical profiles

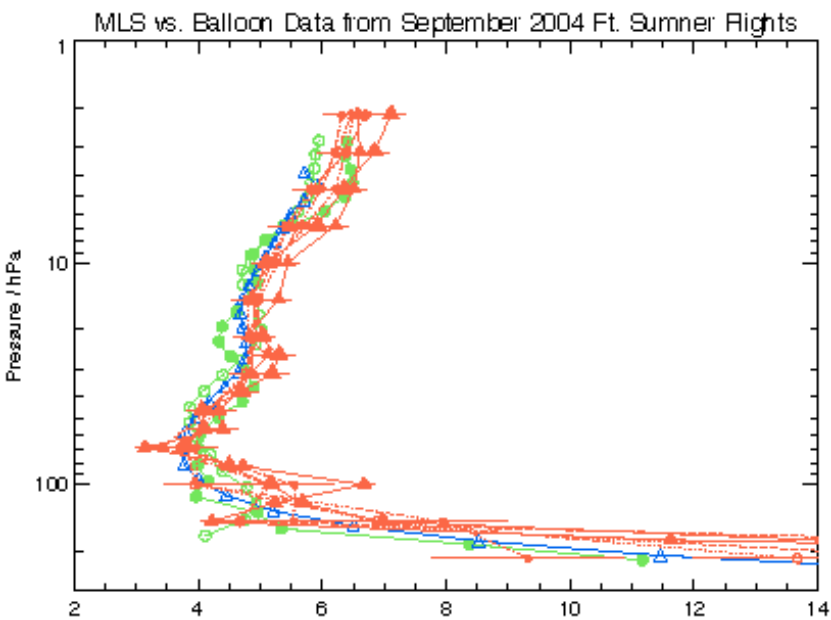
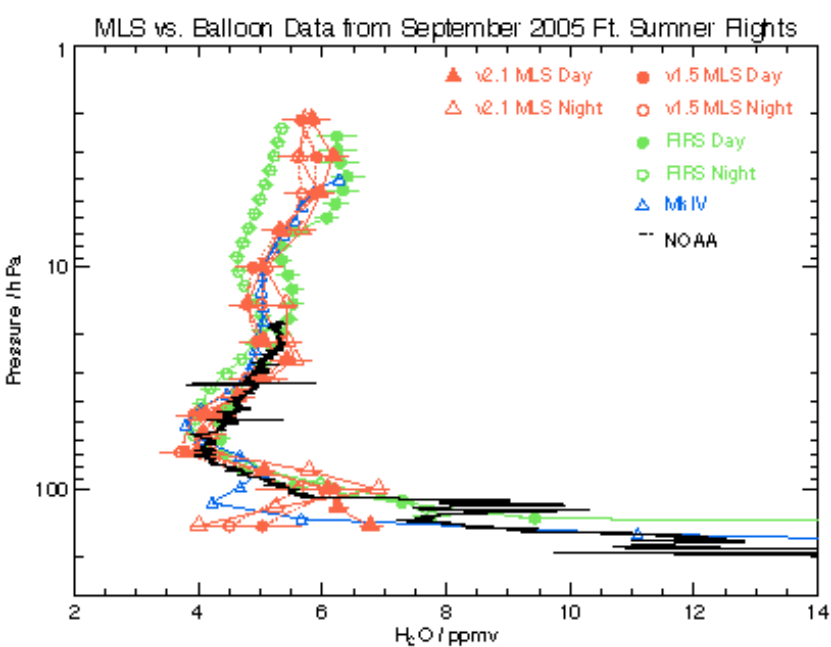
- MLS versions **v2.10** and **v1.51**
- Left: Mean vmr profile
- Right: RMS precision profile

► Lower panel: Global mean difference vertical profiles

- Left: Statistics of profile differences in vmr units
  - Mean difference (v2.10 – v1.51)
  - RMS difference
  - Expected RMS difference
- Right: as Left except shown as a percentage

*Alyn Lambert, H<sub>2</sub>O*





## MLS Strat H<sub>2</sub>O summary

MLS v2.10 stratospheric H<sub>2</sub>O is 0.5 ppmv higher at all altitudes than v1.51

The estimated precision of MLS v2.10 H<sub>2</sub>O is < 0.5 ppmv for pressures > 0.1 hPa

A stratospheric wet bias is seen in the comparisons with ACE, however, there is better agreement with MIPAS in the low stratosphere and at the stratopause than for v1.51 data

Vertical oscillations can be seen in the low stratosphere in single profile comparisons

Further refinements of the MLS Level-2 H<sub>2</sub>O data product will address these issues

Ft Sumner Balloon comparisons  
(Sept 2005 and Sept 2004)

*Alyn Lambert, H<sub>2</sub>O*

# V2.1 MLS H<sub>2</sub>O & RHi Measurements

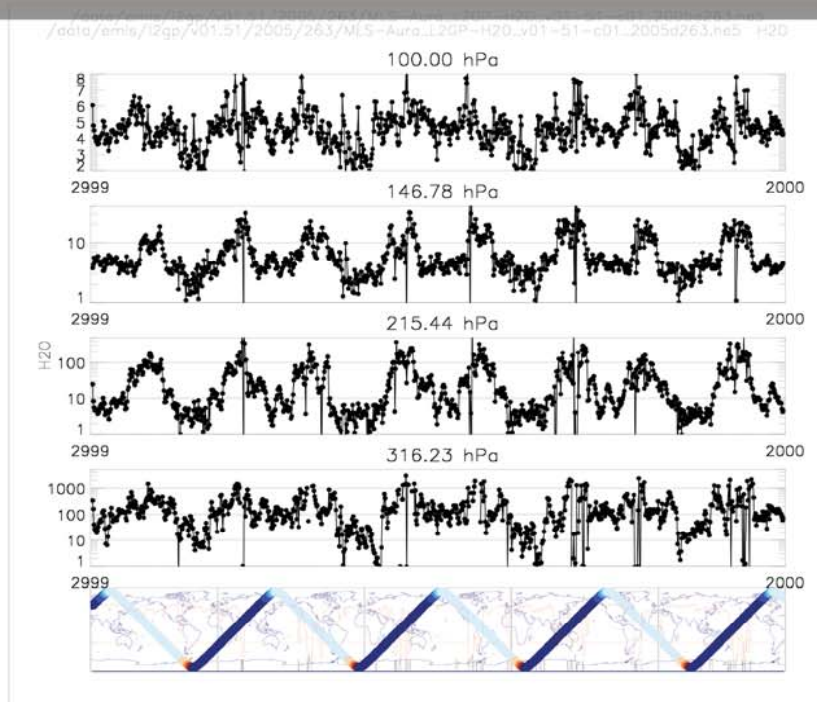
- Coverage is 82°S – 82°N 240 profiles per orbit, ~14.5 orbits per day (~3500 profiles daily).
- Vertical coverage is 681(sometimes)–0.001 hPa.
  - **Relative humidity** is retrieved at 681 and 464 hPa using a saturated radiance slant path technique (similar to nadir sensors).
    - Resolution is 100 km (along track) X 6 km (cross track) X 4 km (vertical).
    - **Not sensitive to temperature errors---good for supersaturation detection but derived specific humidity accuracy strongly depends on T accuracy.**
    - Single profile precision is 40% at 681 hPa and 17% at 464 hPa.
  - **Specific Humidity (or H<sub>2</sub>O)** is retrieved from 316—0.001 hPa using the spectral radiance limb viewing technique.
    - Retrieved every 1.3 km from 316—22 hPa and more coarse above 22 hPa.
    - Resolution is 160 km (along track) X 6 km (cross track) X TBD (~2.5—3 km troposphere).
    - **Not sensitive to T accuracy but derived RH will be.**
    - Single profile precision is ~5% between 316—83 hPa.
  - Will focus on the 681, 464, 316, 261, 215, 147, 121, 100 and 83 hPa levels here and will only show specific humidity.

18 days of V2.1 H<sub>2</sub>O currently available

*Bill Read*

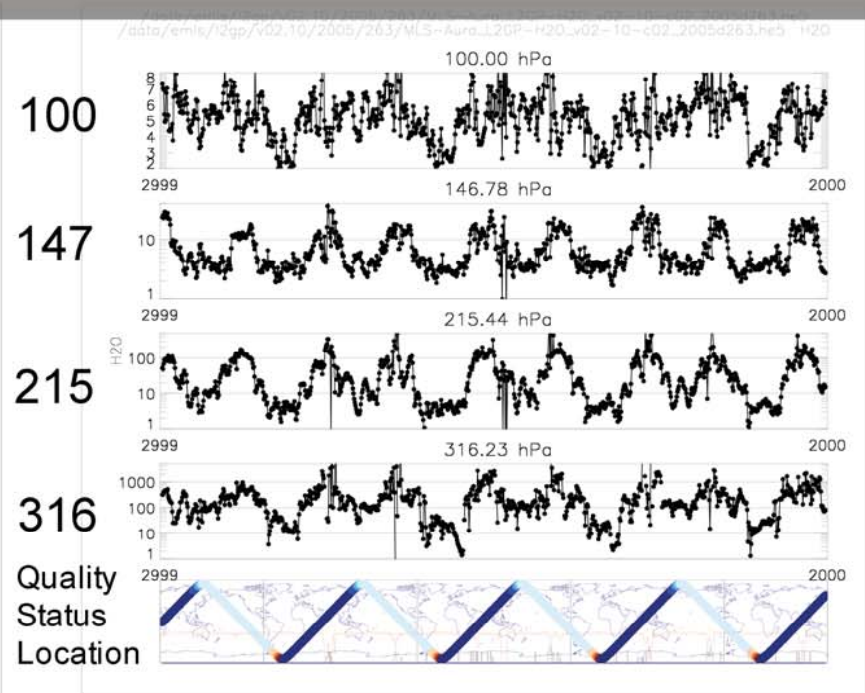


## MLS v1.51



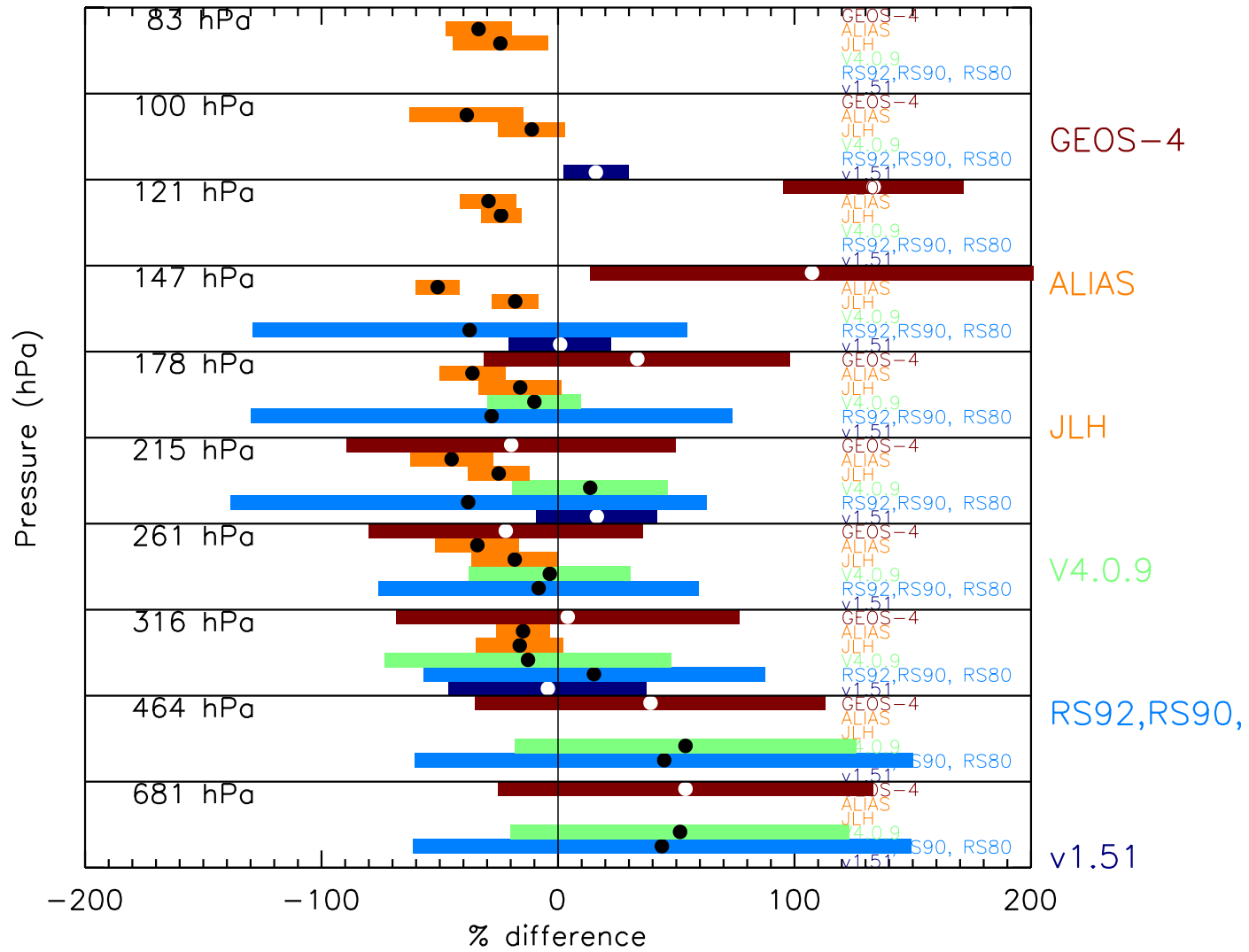
- MLS v1.5 had problems with dry spikes at 316 hPa.
- Poor convergence in clouds.
- 10—25% precision in troposphere.

## MLS v2.1



- Time series look similar.
- Note less noisy appearance of 215 and 316 levels. Fewer dry spikes.
- 100 hPa is more ragged.
- Poor convergence in clouds.

# Summary



# Conclusions

- Mean differences with Vaisala (RS80/90/92) radiosondes between 316—261 hPa and with AIRS between 316—177hPa are < 20%.
  - The scatter about the mean is large, typically about 60-70% for the radiosondes and AIRS at 316 hPa.
  - With AIRS, the scatter is drops to ~30-35% between 261—177 hPa.
  - Important not to compare AIRS humidity measurements when  $H_2O < 50$  ppmv.
- MLS is consistently 10—25% drier than JLH at all altitudes. ALIAS is very close to JLH except when it encounters cirrus. On the 22 Jan 2006 flight, a cirrus event was detected (large difference between JLH and ALIAS) which did not appear to impact the MLS  $H_2O$  measurement.
  - Although the spatial and temporal coverage of the in-situ is much more limited, the scatter is smaller than the best achieved with other satellite— even down to 316 hPa.
  - Perhaps benefiting from comparing like measurements (specific humidity).
- At this time the MLS relative humidity measurements at 681 and 464 hPa do not appear to be of high quality.
  - They are biased high.
  - Large scatter about the mean.
- GEOS-4 shows larger scatter with MLS than with other measurements.
  - GEOS could benefit from assimilating the MLS humidity.

Looking toward v2.2

- Improve the 681—464hPa RH retrieval.
- Improve convergence: currently ~60%.
- Less noisy at 100 & 83 hPa.

*Bill Read*

# HIRDLS

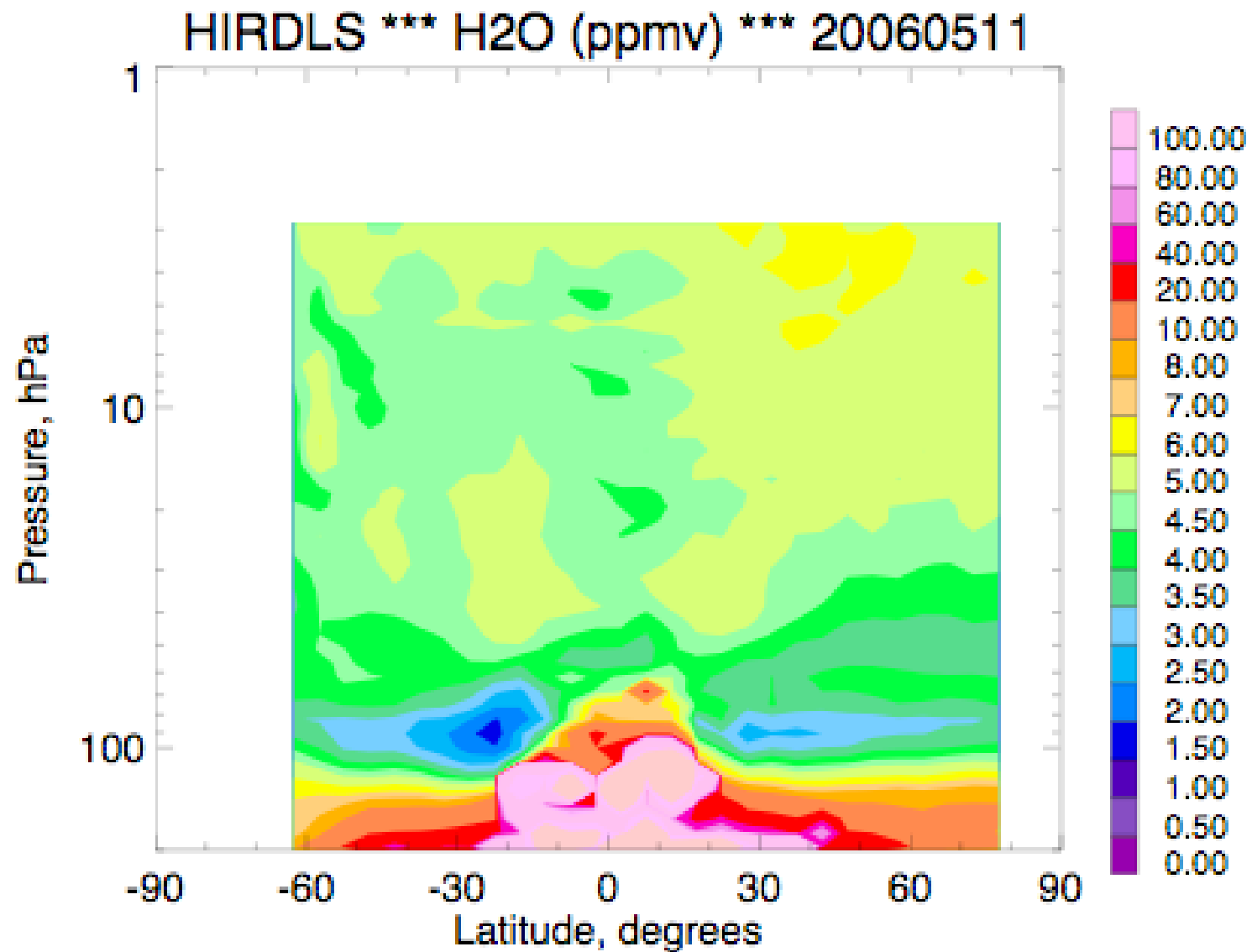
## **Caveats:**

**Water vapor results very sensitive to oscillation perturbations**

**Most effective versions of the “Deoscillation” algorithms are very new**

**Useful water vapor results are also quite new, so not much time to study in detail. These are first looks.**

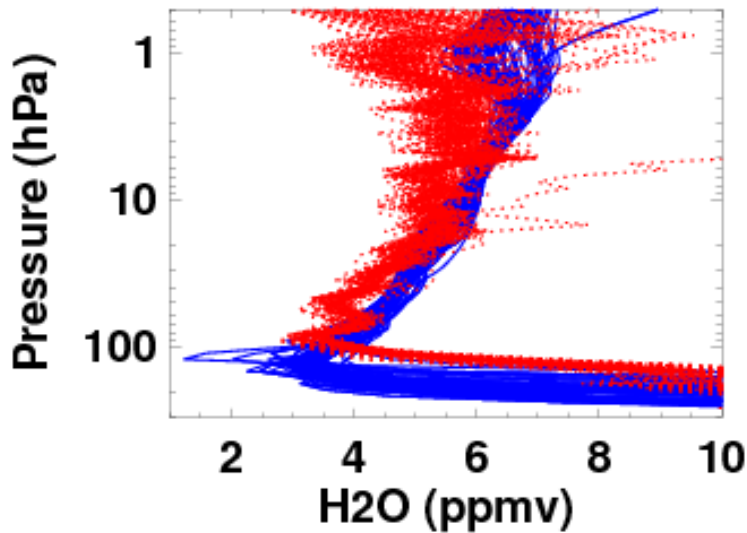
# Zonal Mean Water Vapor





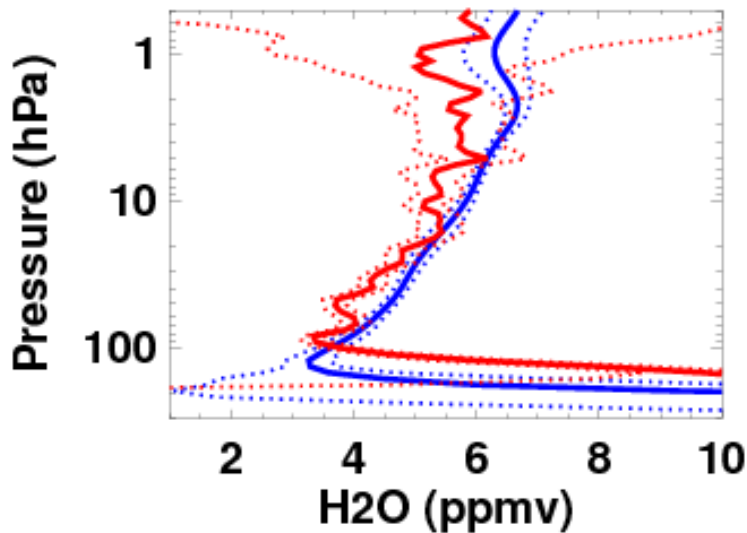
# **HIRDLS & ACE Water Vapor Profiles**

**NH H<sub>2</sub>O**



**All Coincidences  
Within 2 hours**

**NH H<sub>2</sub>O**

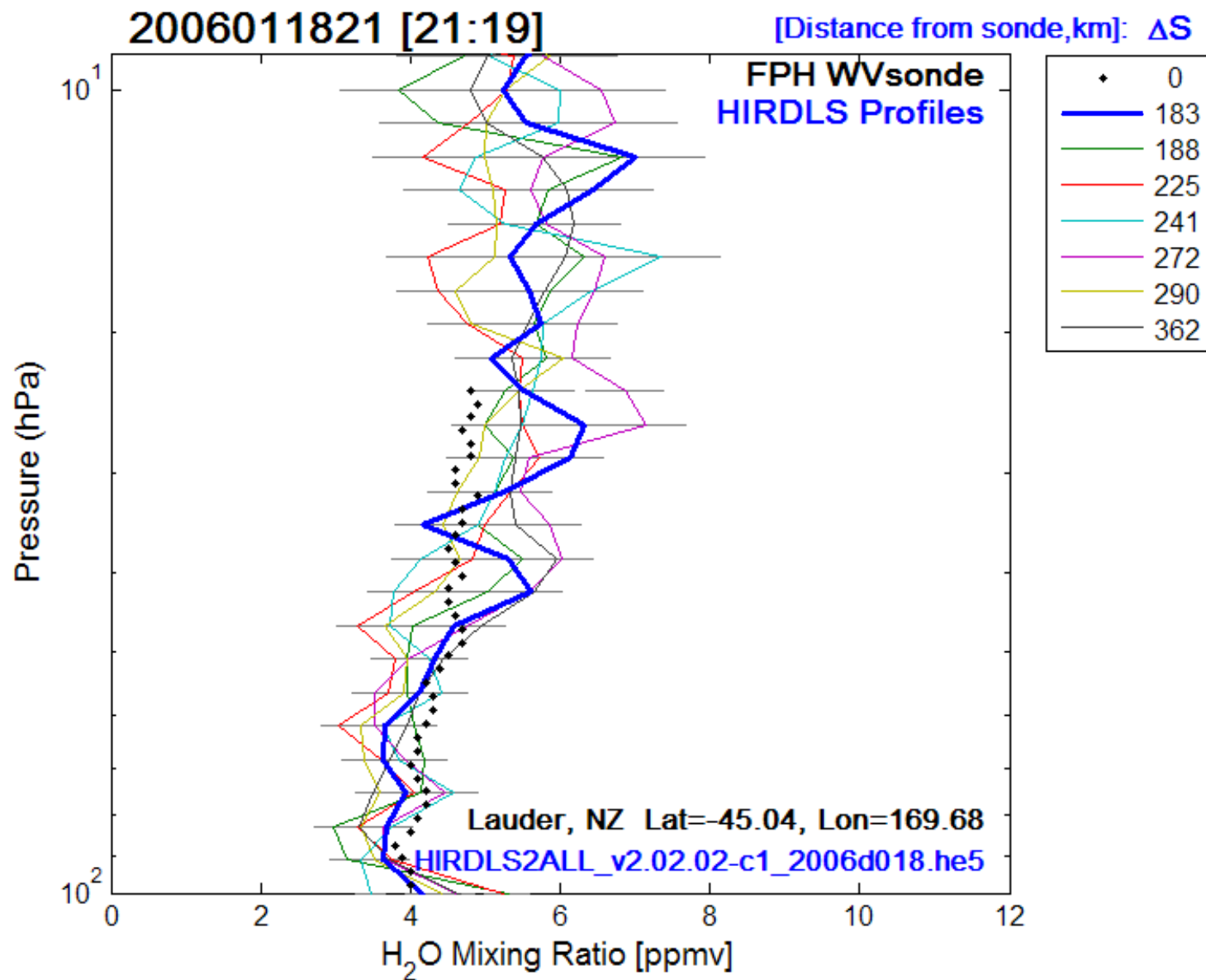


**Average (solid) & 1- $\sigma$   
standard deviation  
(dotted)**

Cora Randall, Peter Bernath and the ACE Team

*John Gille*

# Sonde Comparison- Lauder NZ



# Summary

**Water vapor cross-section and zonal means have reasonable values**

**Some evidence of residual oscillation for some scan tables (refine)**

**Values too high in tropics above the tropopause (blockage correction)**

**Problems-**

**Small scale horizontal variability**

**Small scale vertical variability**

**Possible problems at high latitude, high altitude**

**Data are clearly on the right track, much further refinement is needed.**

# **Intercomparisons of the Harvard Lyman alpha hygrometer and ICOS isotopic water instrument with the CFH and MLS instruments: Implications of recent results**

## *Questions to be explored*

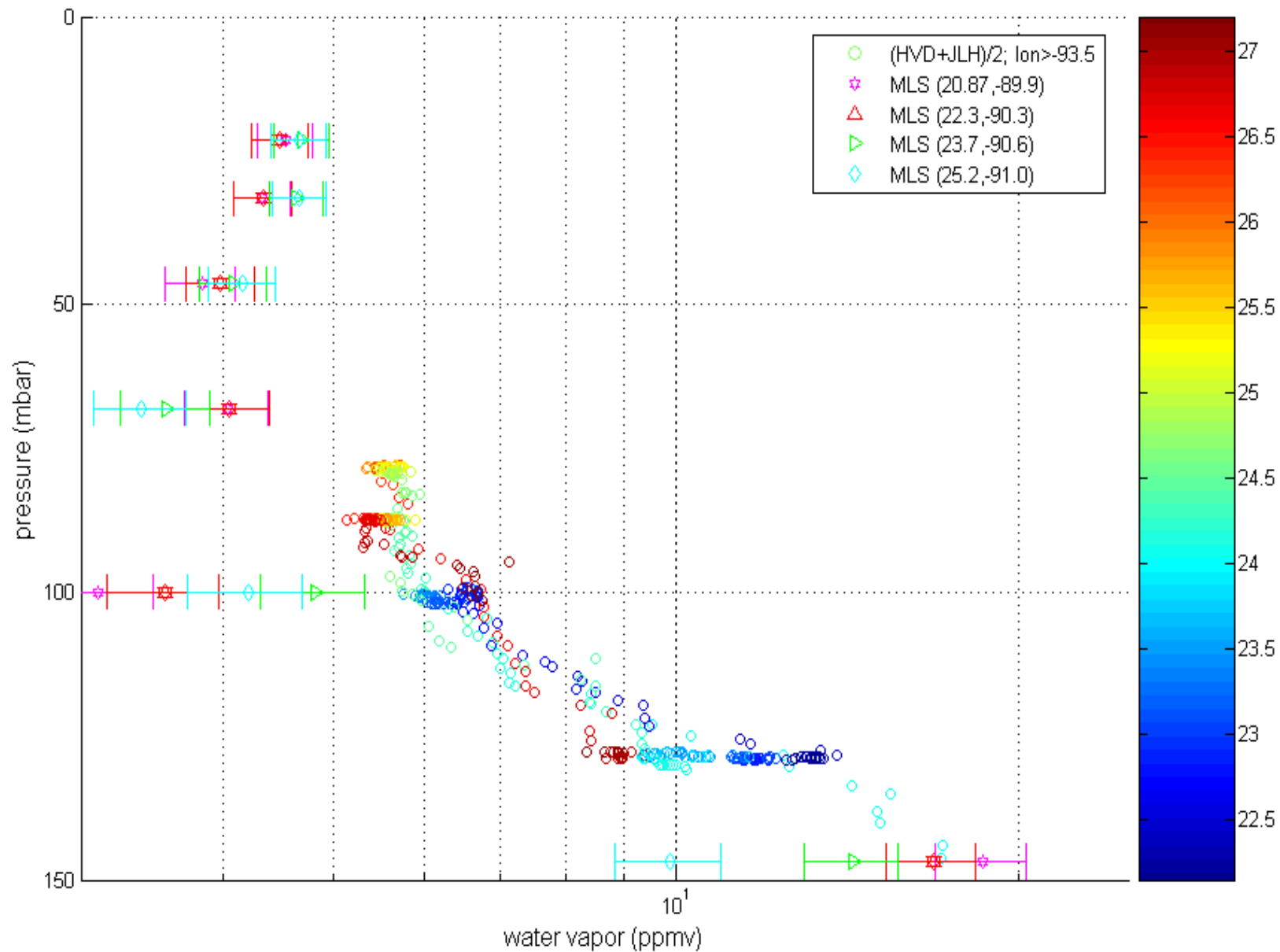
Are intercomparison data from CRAVE and AVE-WIIF consistent?

What have we learned from CRAVE regarding the accuracy of in situ water instruments needed for Aura satellite validation, especially regarding the previously observed systematic differences between the frost point hygrometer and in situ aircraft instruments?

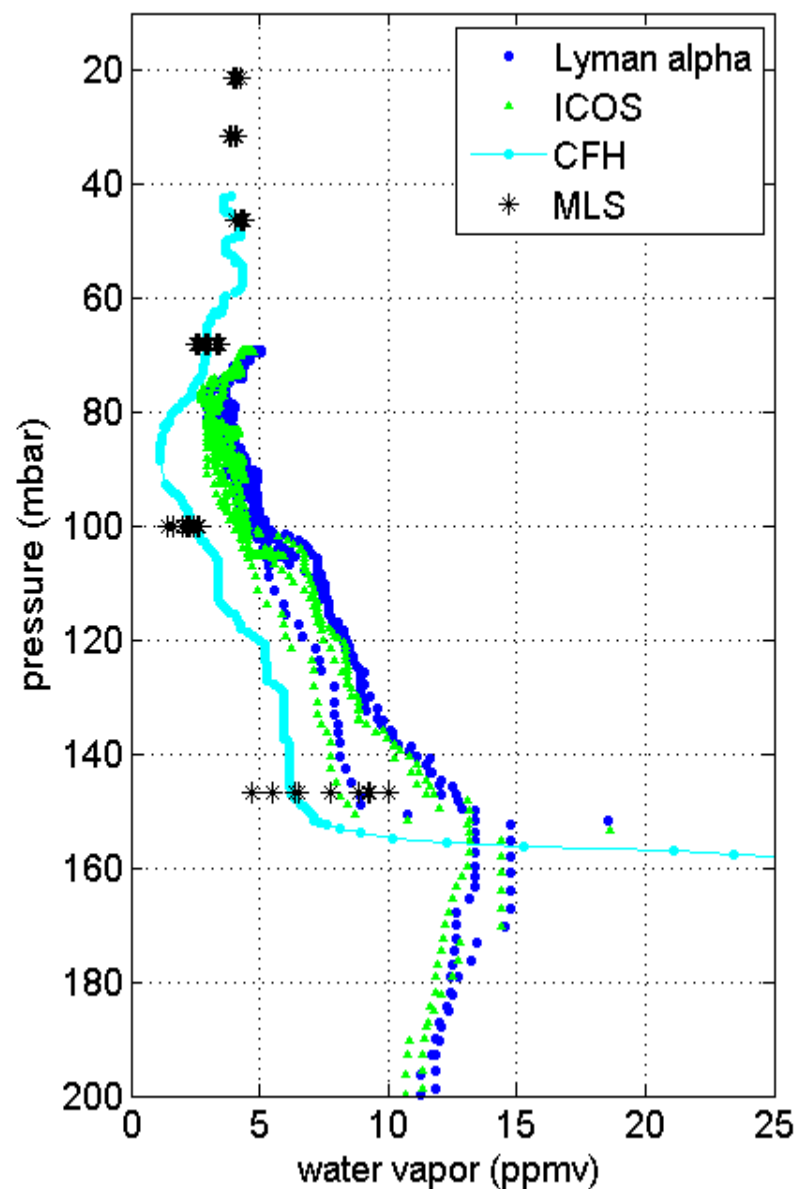
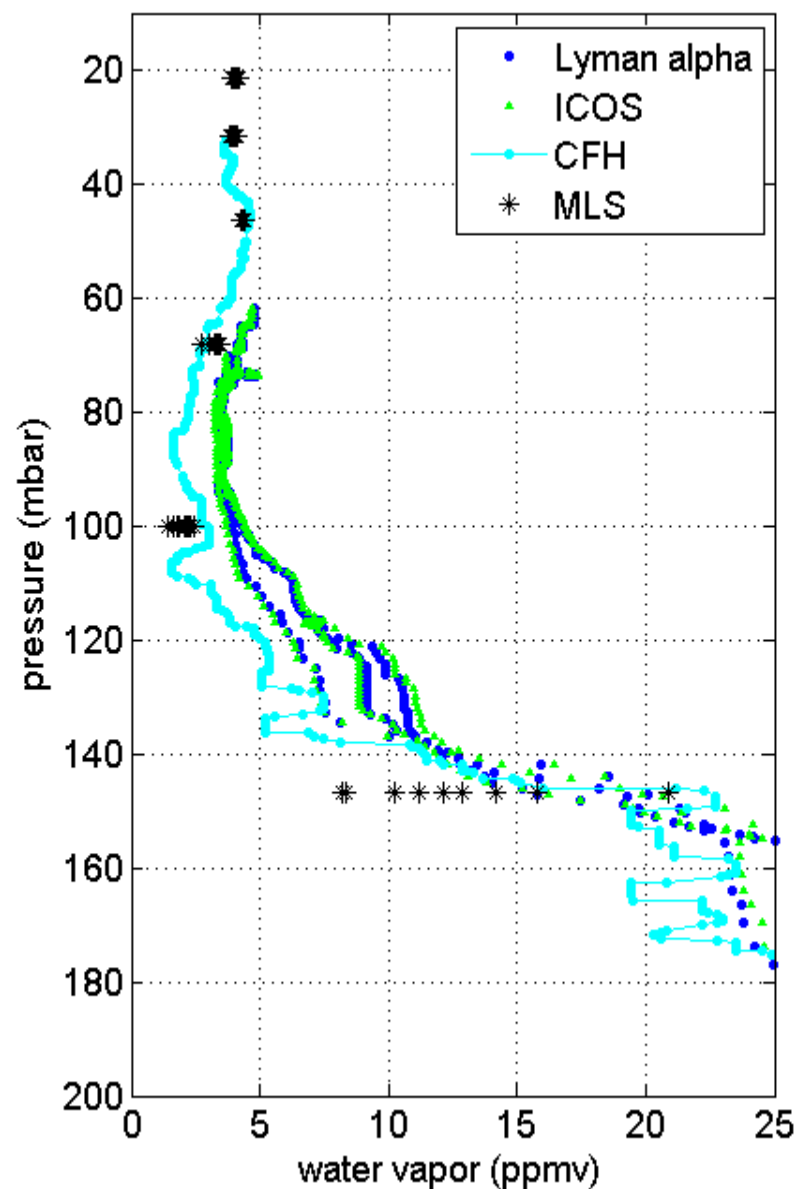
How do MLS version 1.5 and version 2 compare with in situ water vapor measurements?

*Elliot Weinstock*

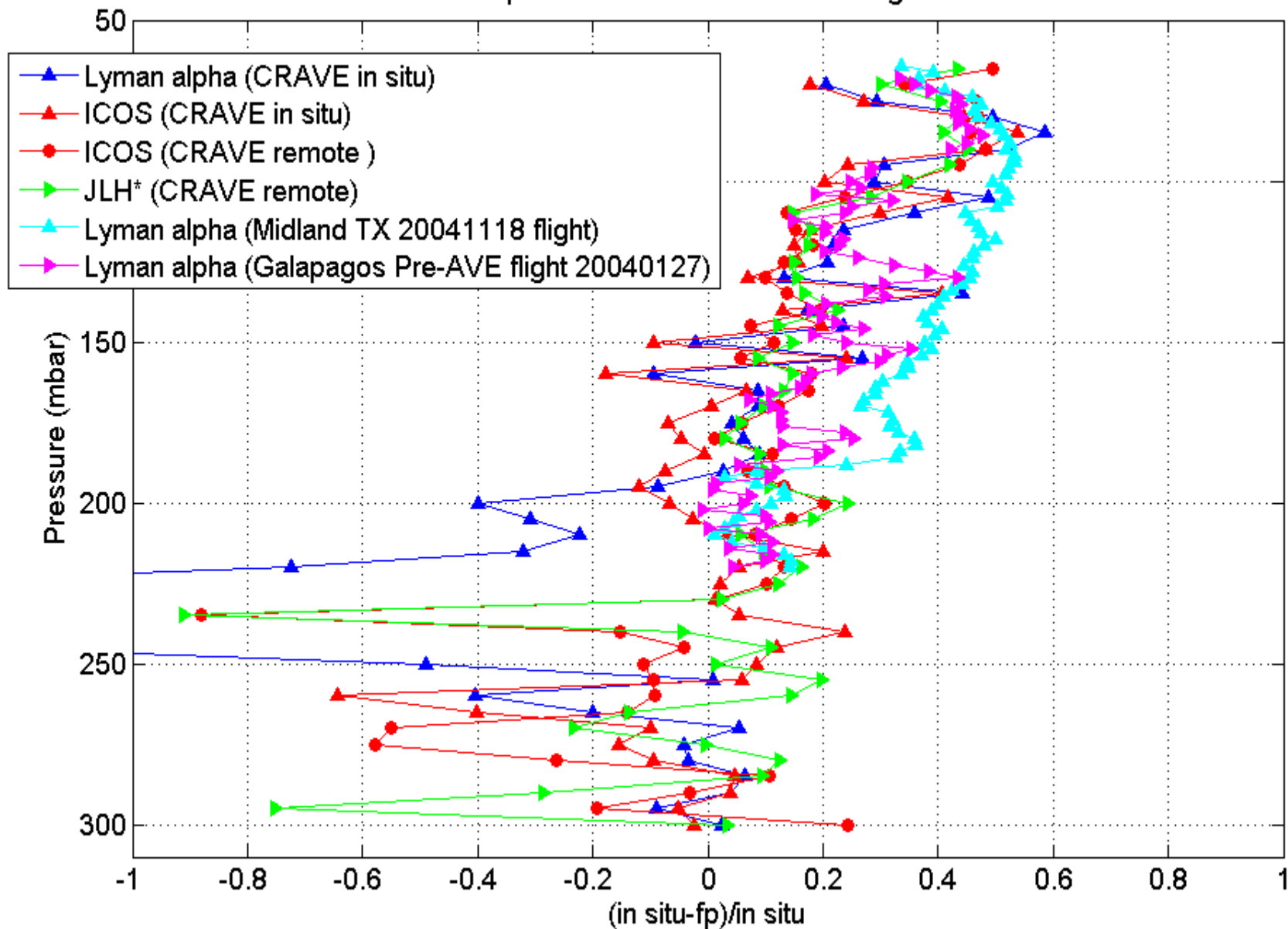
# AVE-WIIF 20050703



*Elliot Weinstock*

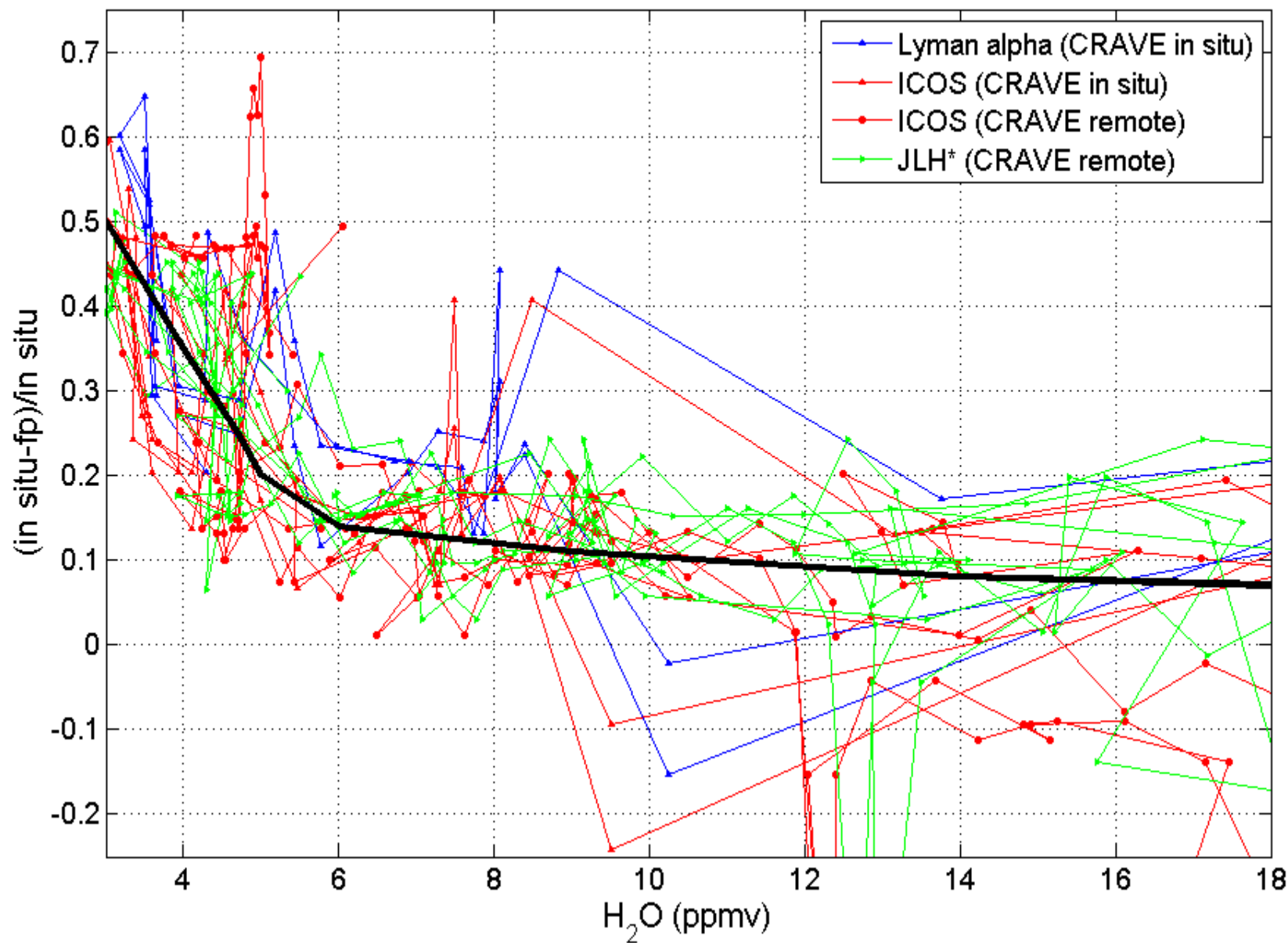


# Intercomparison of in situ and CFH during CRAVE



*Elliot Weinstock*

# Intercomparison of in situ and CFH during CRAVE



*Elliot Weinstock*

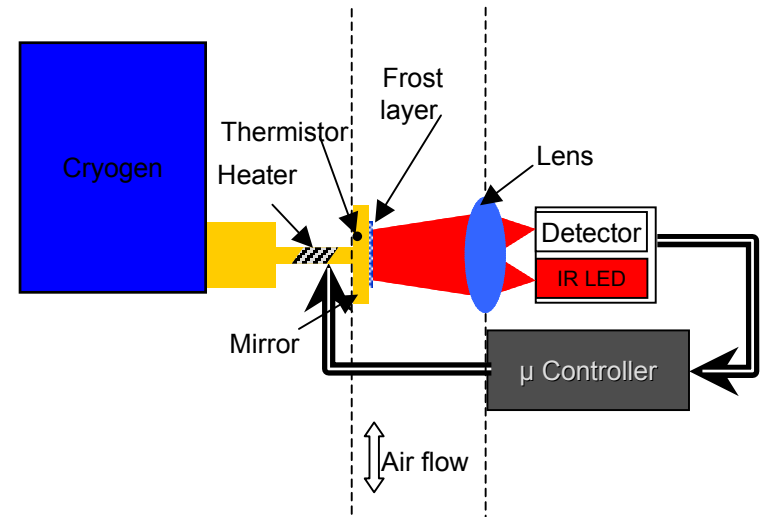


# Conclusions

- As in AVE-WIIF, the overall agreement between Harvard water vapor instruments during CRAVE was very good.
- Comparisons between in situ water vapor on the WB57 and the CFH instrument illustrate systematic differences that increase significantly at low water vapor.
- Missions that provide the opportunity for careful water intercomparisons continue to be very useful and need to continue.
- Laboratory intercomparisons with low water vapor mixing ratios need to be carried out to help determine the source of this discrepancy.

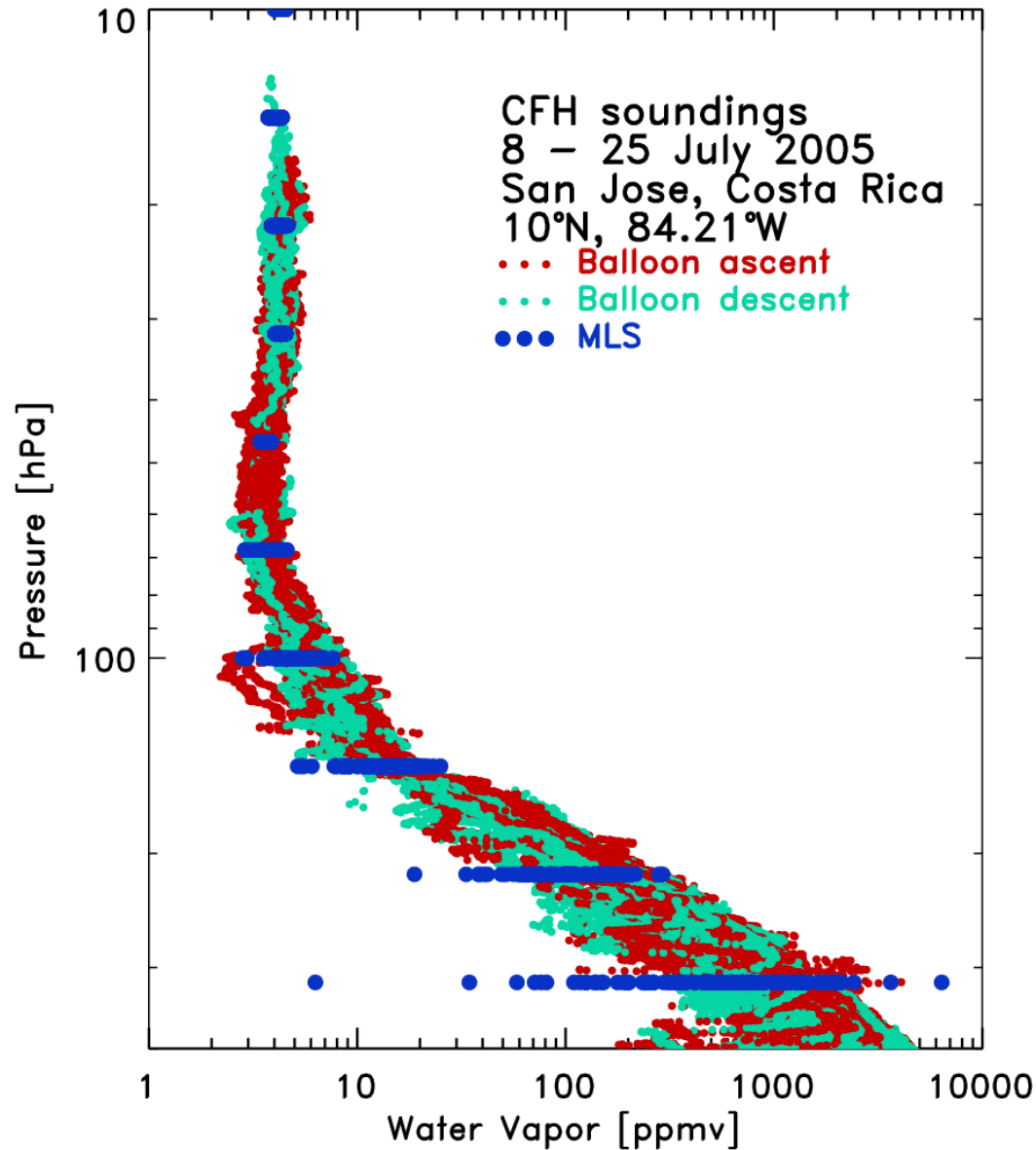
# Cryogenic Frostpoint Hygrometer (CFH)

- Absolute measurement
- Vertical Range: surface to ~28 km  
(surface to ~25 km on ascent)
- Uncertainty: troposphere: > 4% MR  
stratosphere: ~ 9 %
- Microprocessor control
- Phase sensitive detector:  
electronic sunlight filter
- Weight: ~ 400 gr
- Payloads carry ECC ozone  
sonde and Vaisala RS80
- ~170 soundings so far



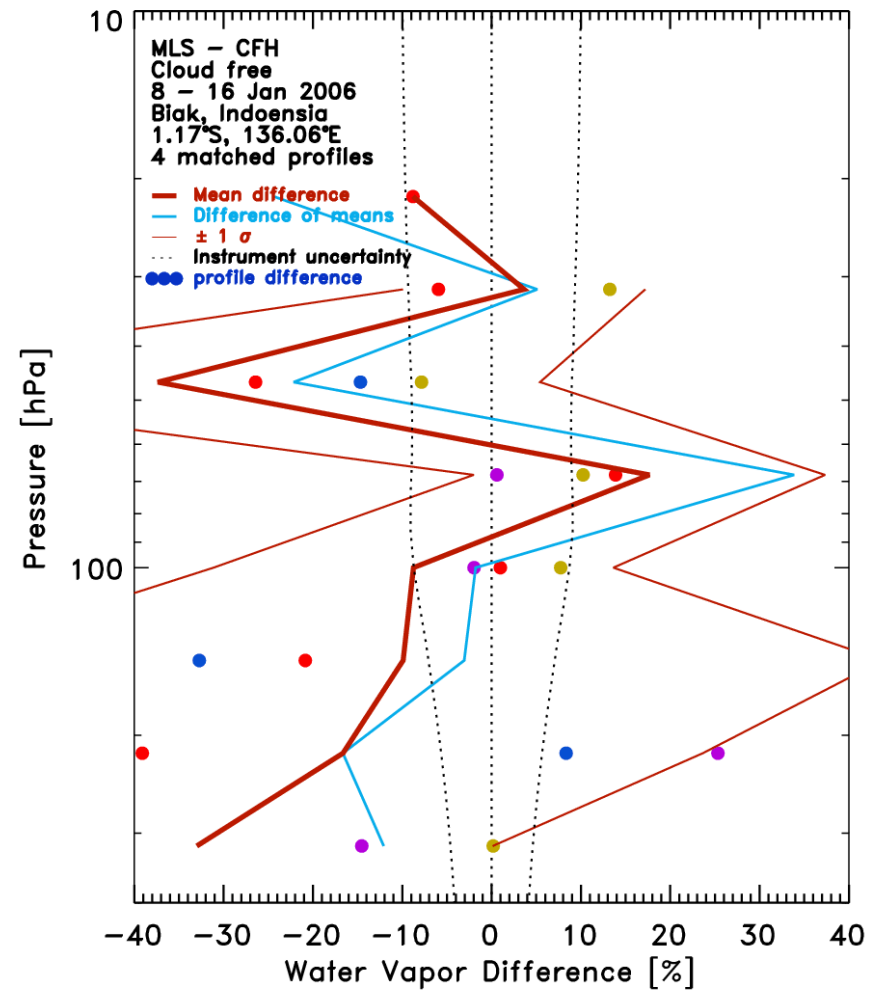
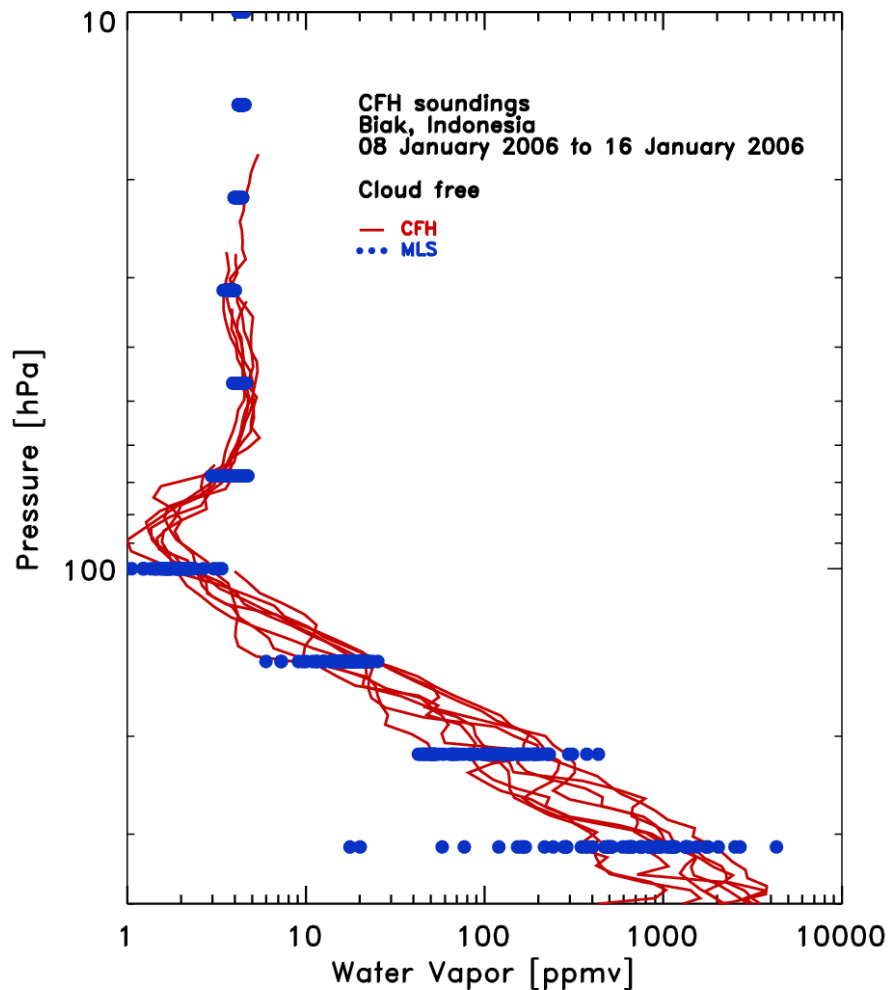
*Holger Vömel*

# Satellite comparison



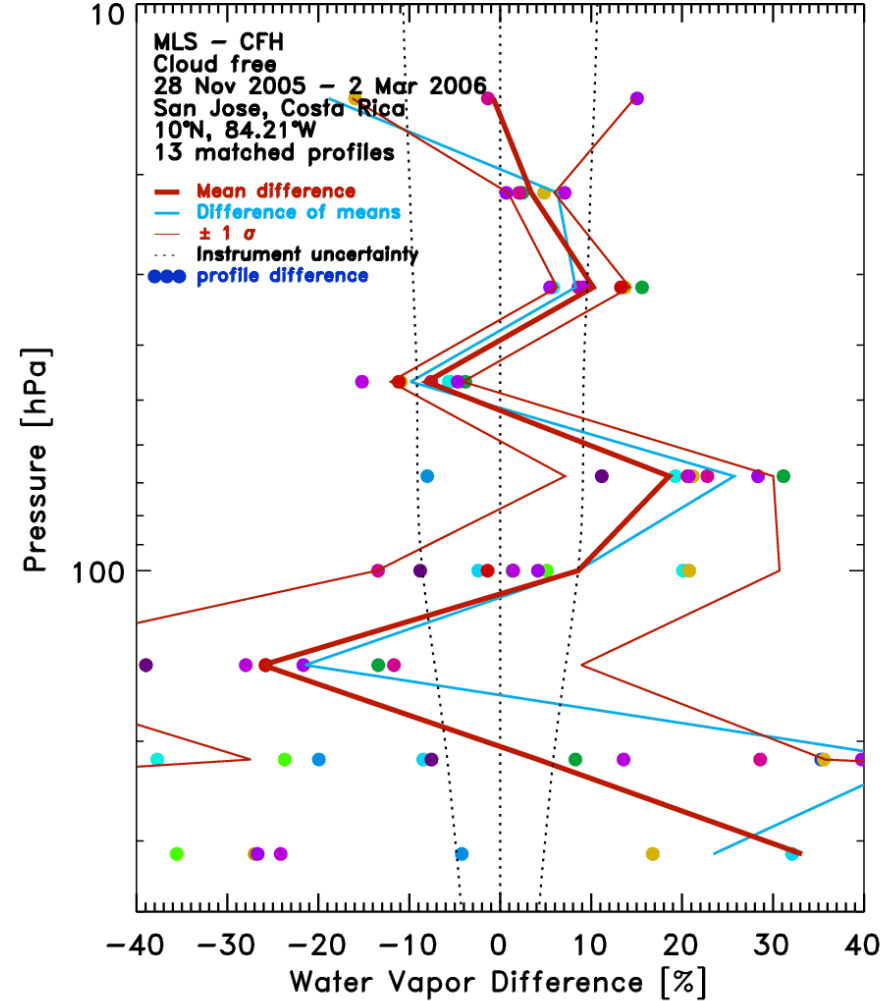
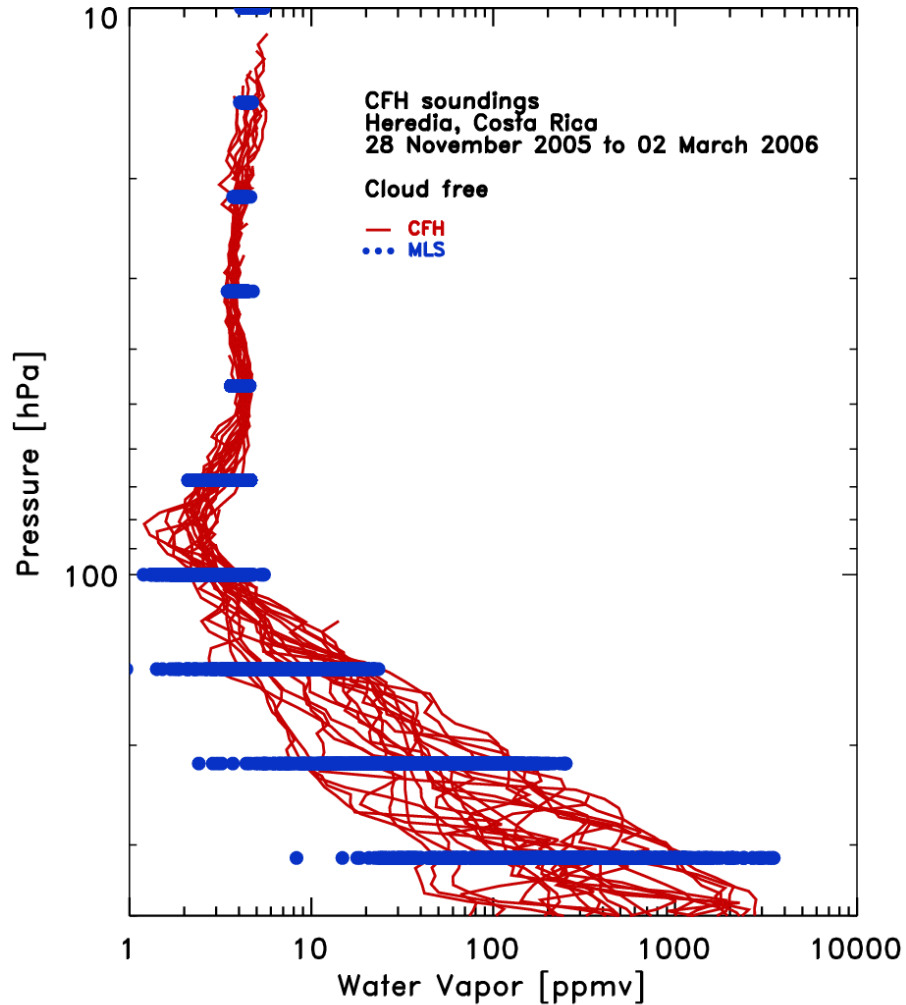
# MLS Comparison: Tropics

*Biak Indonesia: Jan 2006*



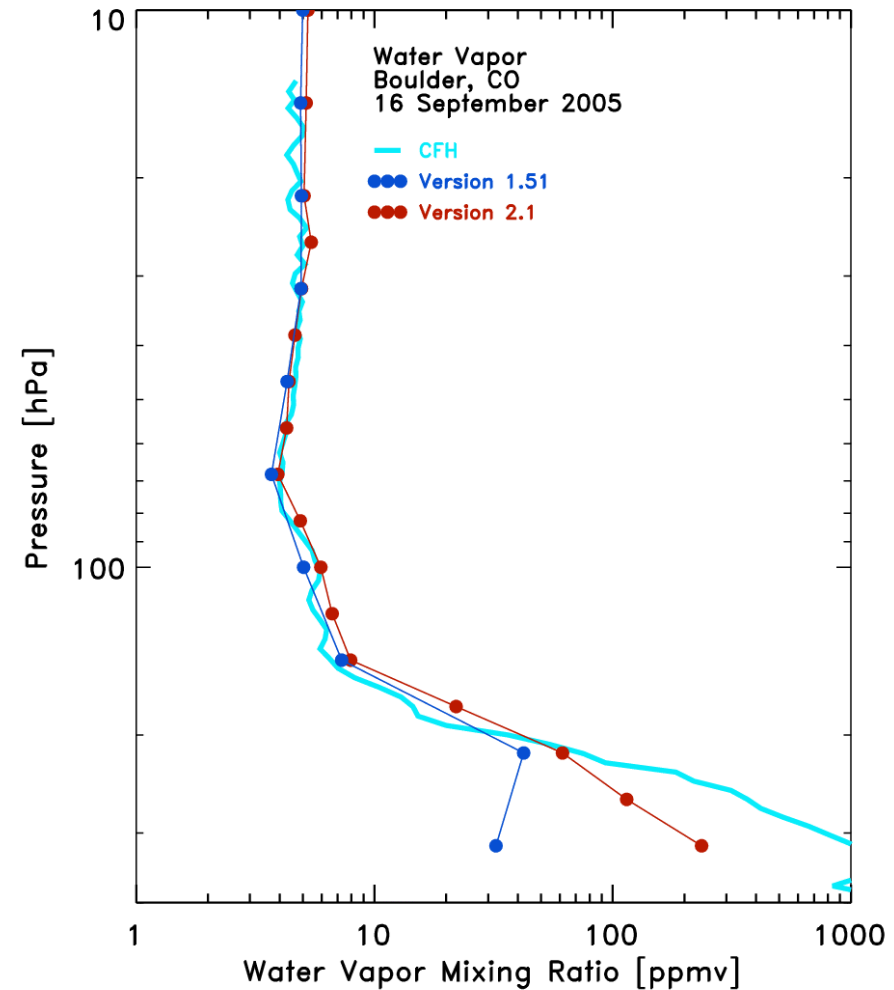
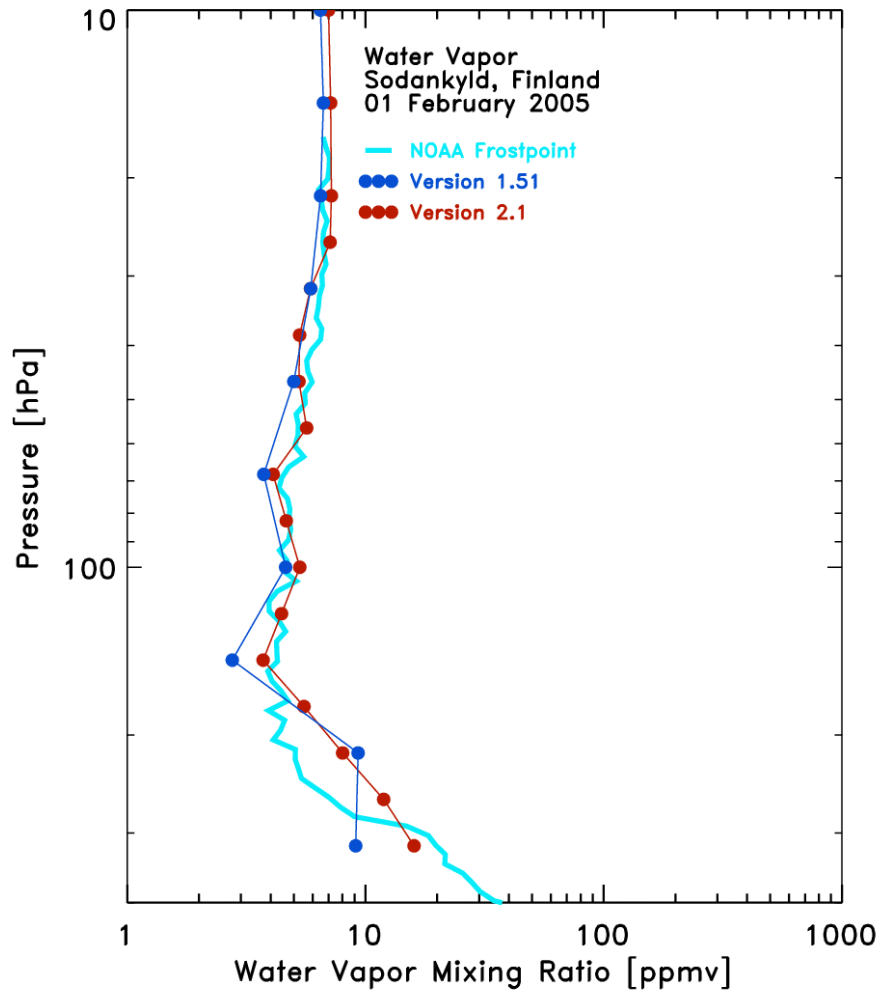
# MLS Comparison: Tropics

*Costa Rica AVE Jan/Feb 2006*



*Holger Vömel*

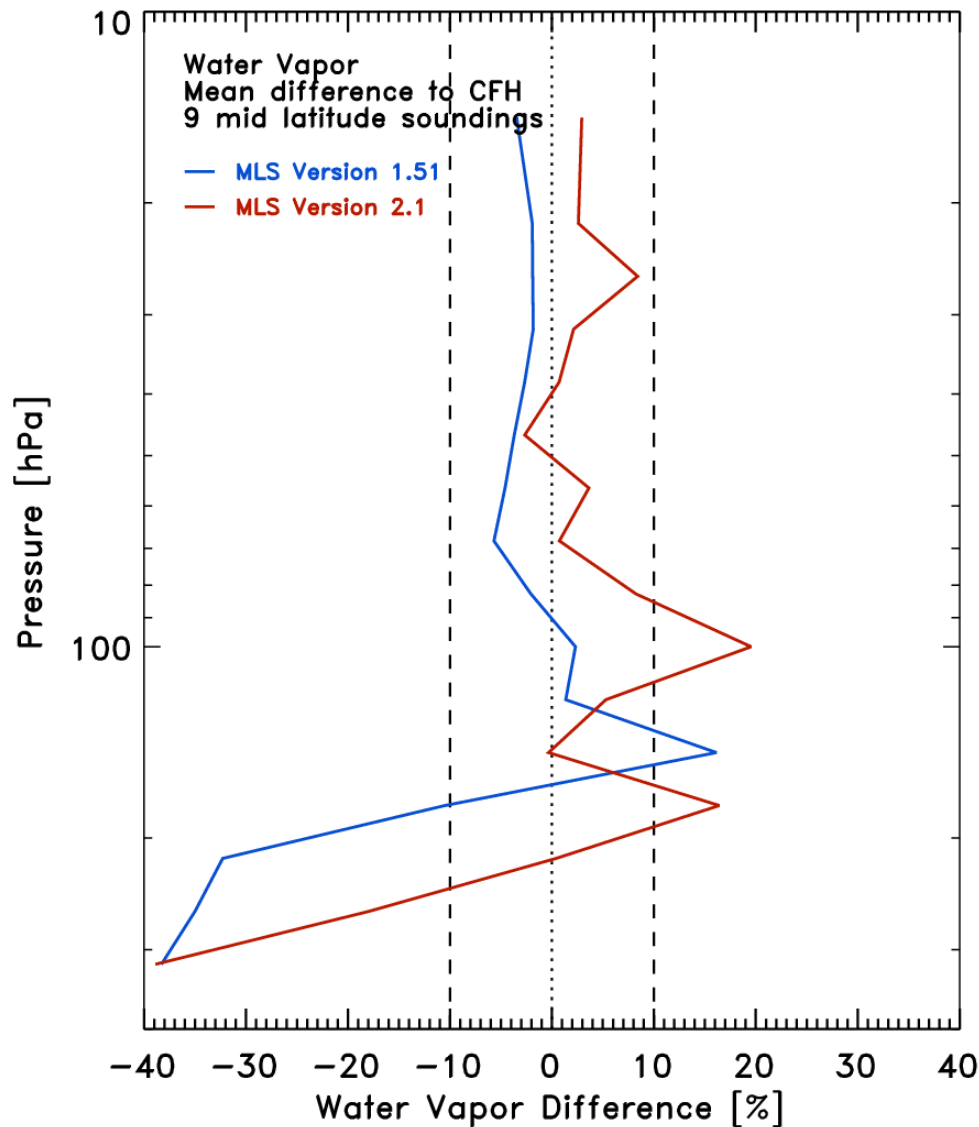
# Sondakylä & Boulder example



1.51 and 2.1 and CFH

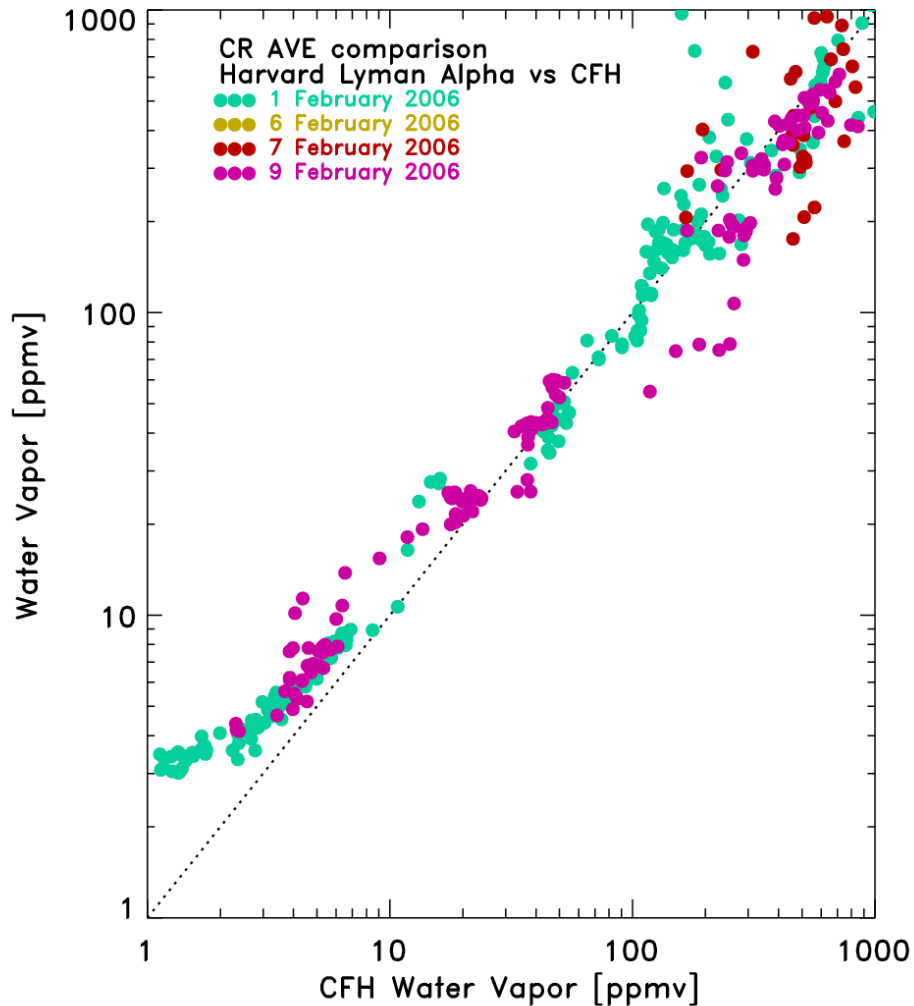
*Holger Vömel*

# Average difference from CFH: version 1.5 & 2.1

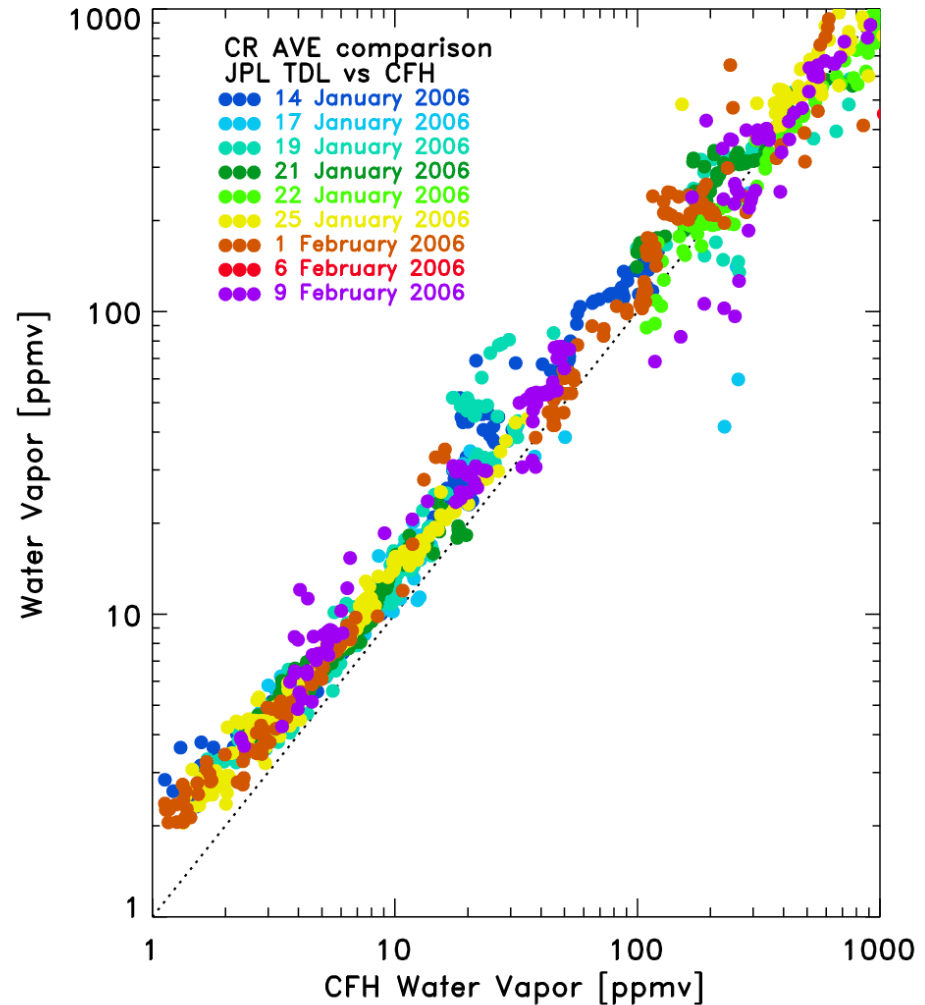


Mid and high latitudes  
only

# CFH Correlations



Harvard Ly-alpha



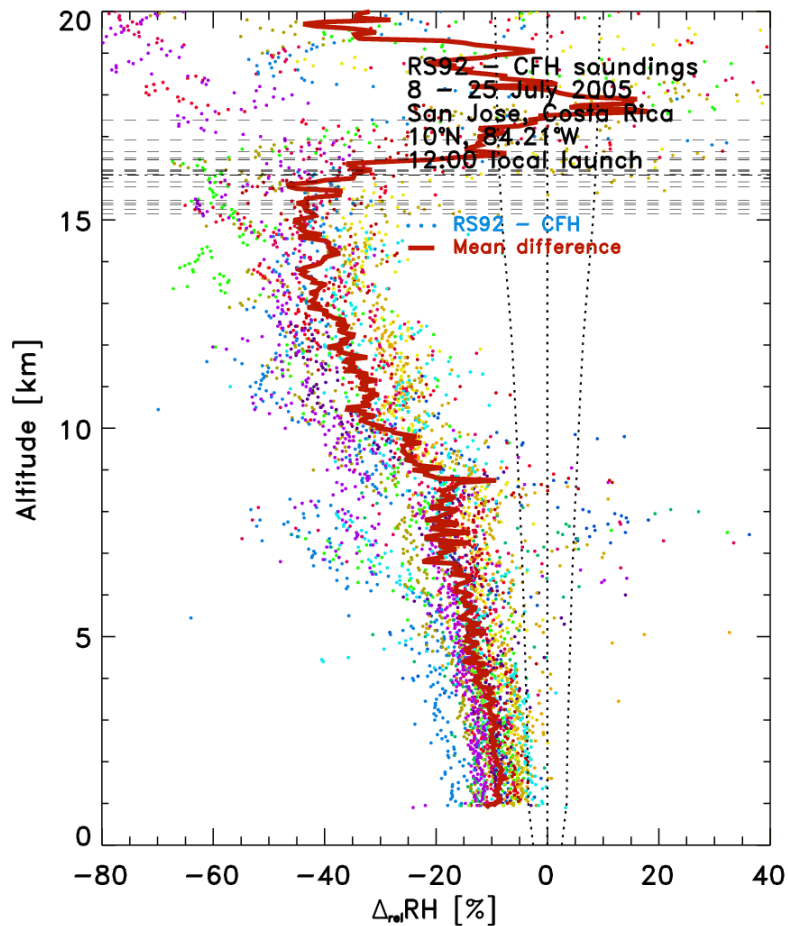
JPL TDL

Constant offset, ly-alpha, scaling factor with JLH

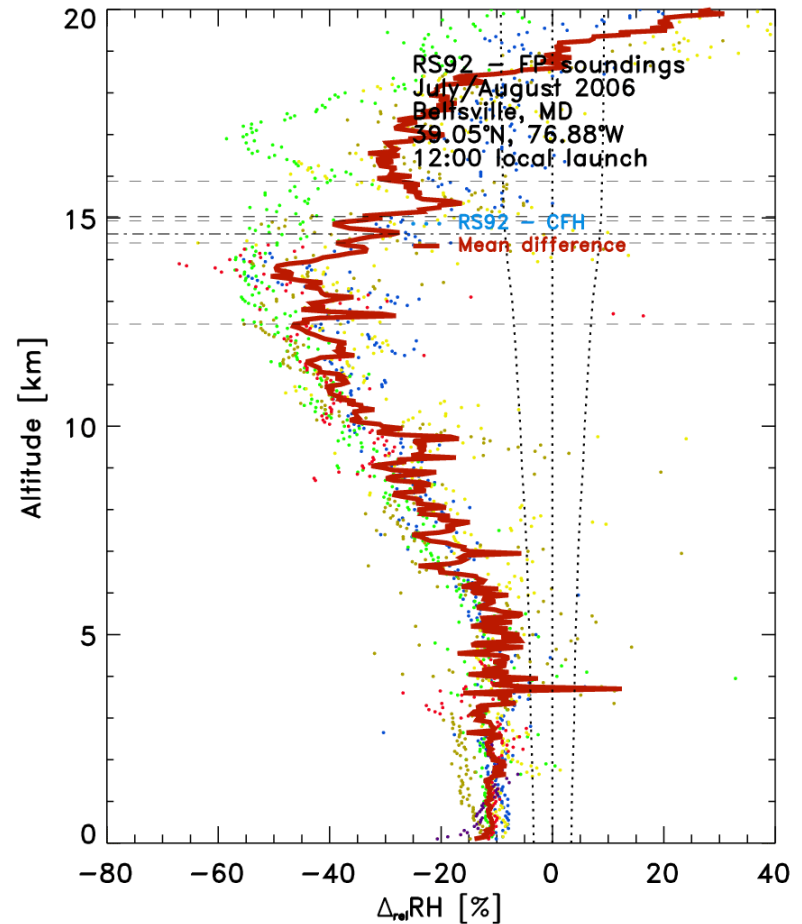
*Holger Vömel*



# Relative RH difference RS92 - CFH



Costa Rica Jul 2005



Beltsville, MD Jul 2006

Shows daytime dry bias of RS92 in UT

# Summary

- Stratospheric MLS water vapor:  
Agreement within measurement uncertainty for both version 1.5 and 2.1
- Except for tropical tape recorder during boreal winter
- Tropospheric MLS water vapor:  
Version 2.1 improves general shape in UT,  
but still very dry and highly variable
- WB57 instruments are too wet compared to CFH;  
no serious disagreements between CFH and other  
balloon or Geophysica instruments
- Vaisala RS92 relative humidity during daytime still up to  
50% too dry (same as last year)

What are the major validation issues that remain?

How does reprocessing affect validation plans?

What additional correlative measurements are needed?

What additional analyses are needed?

What papers are planned/completed at this point?

**One big issue...sorting out why different *in situ* differences at low mixing ratios.**

**In regards to validation papers, the question came up as to whether there will be enough reprocessing done (for MLS).**

**Prioritizing reprocessing to match where correlative measurements exist is needed.**

**TES likely needs more accurate UT measurements (considering problem with daytime operational sonde measurements in the UT.**

**Need to consider continuation of trends in the stratosphere, so matching up existing Aura measurements with past satellite measurements is important (see poster by Brad Sandor).**